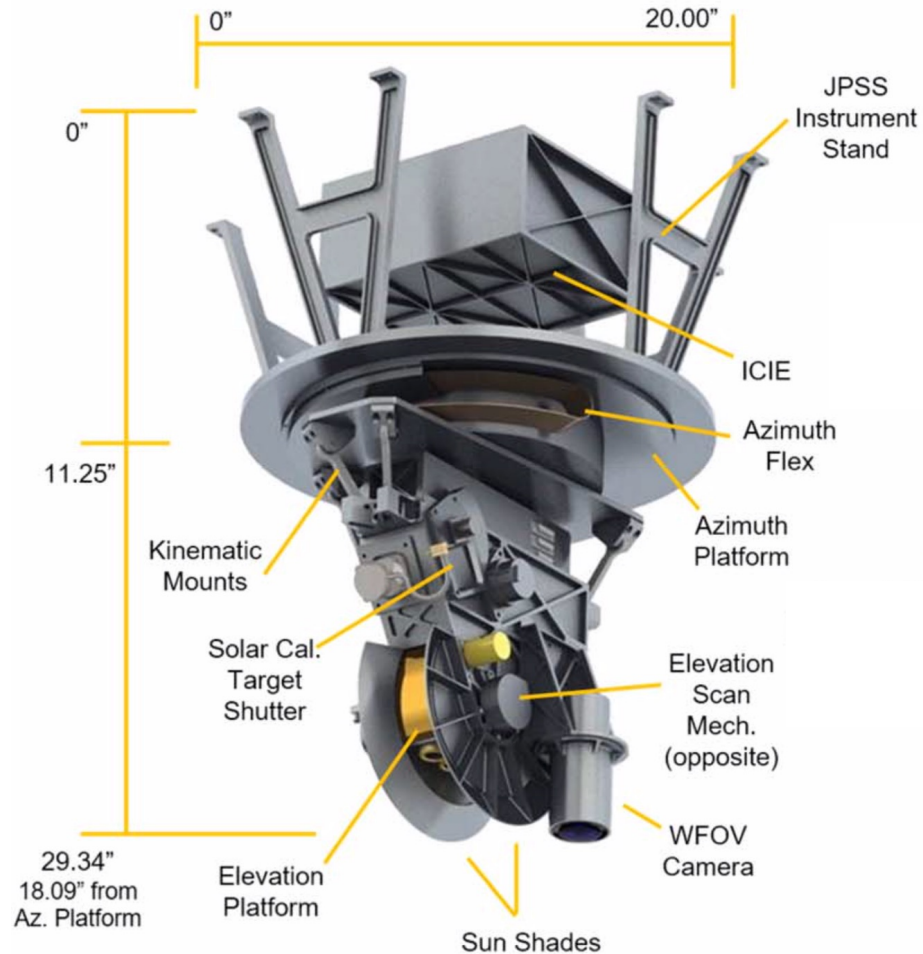


Motivation, characteristics, and use of the Libera wide field of view camera



Libera

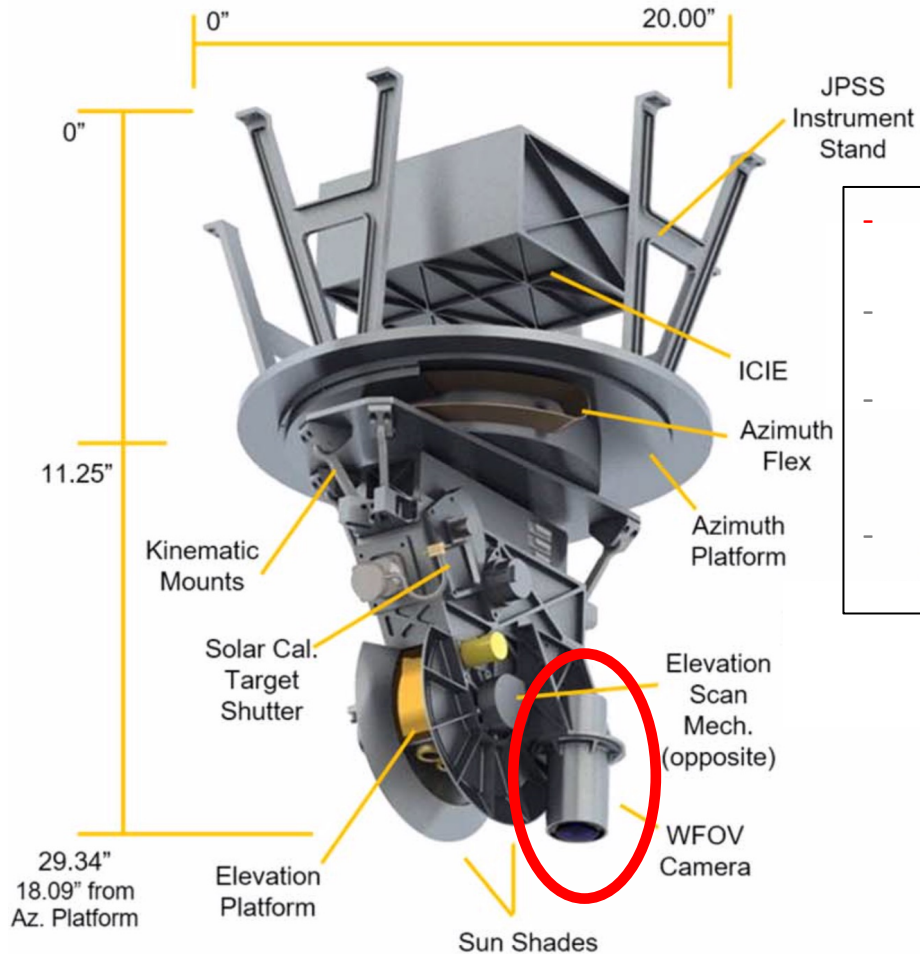


Understanding Earth's Energy Budget

LASP • JPL • LBL • UA • CSU • UM • NIST • NOAA • Ball • SDL

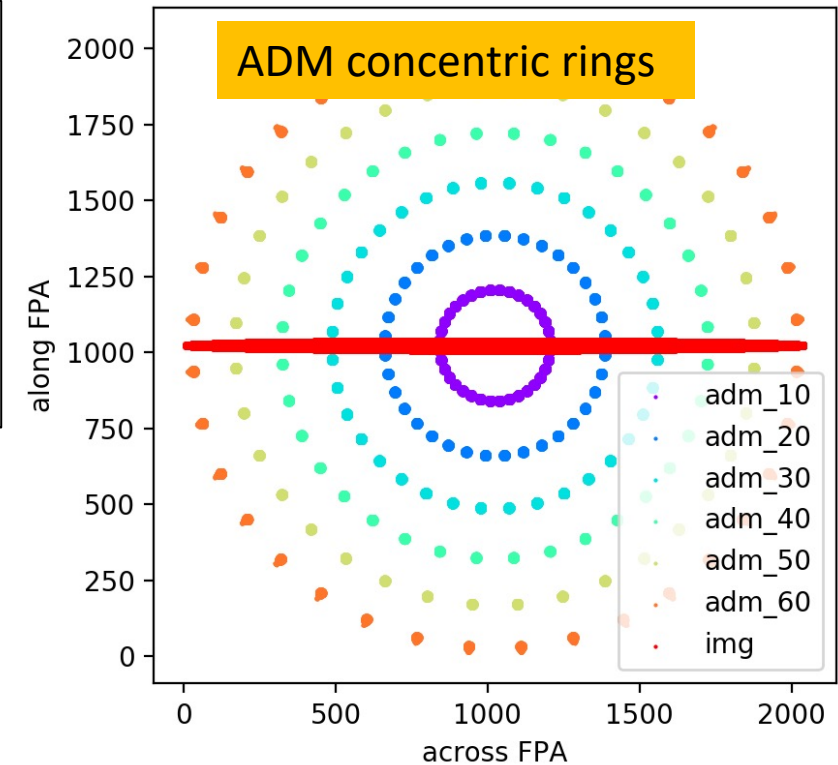
Sebastian Schmidt and the LASP/Ball camera working group

Motivation

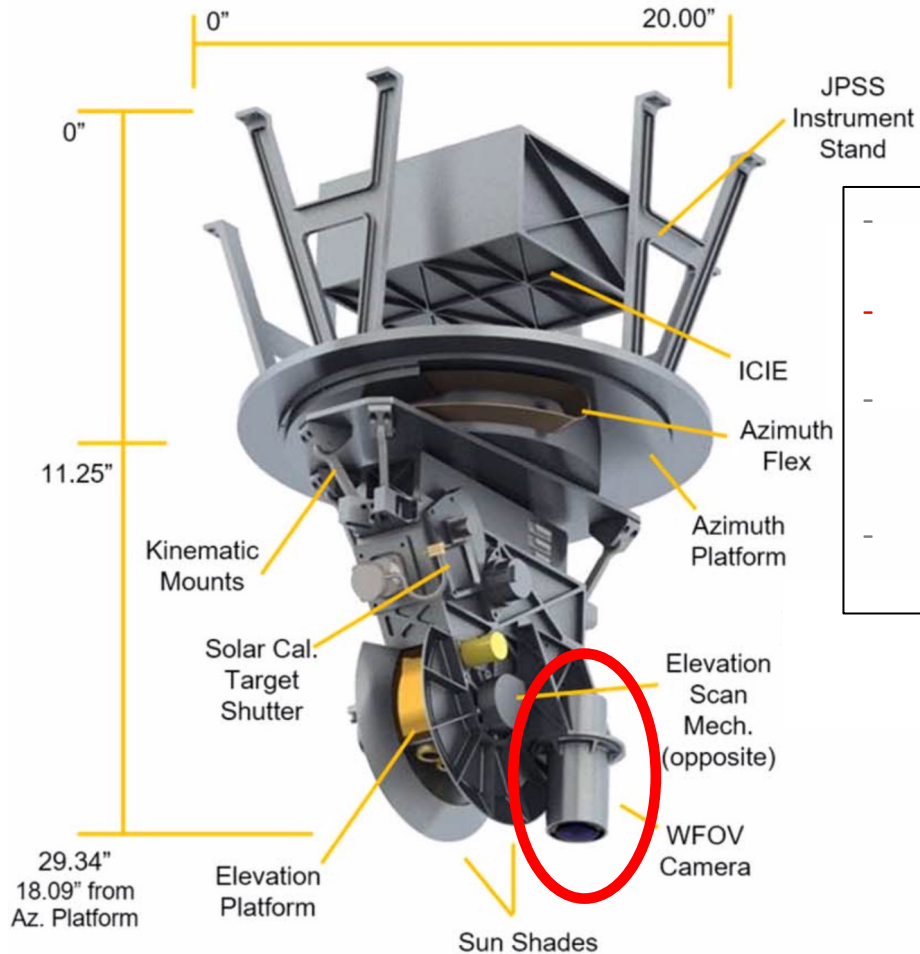


1. Need Angular Distribution Model (ADM) for split solar channel
2. Proof-of-concept scene-type for future separation from imager

- Views multiple angles with the **same imager that is also used for scene type ID**
- Will not replace VIIRS-based scene type ID, but will explore new future avenues
- *Cross-track*: co-alignment with Libera radiometers; used to "fill in" *along-track* direction even while not in RAP mode
- Why single channel? Low-cost demo; emphasis on limb-to-limb

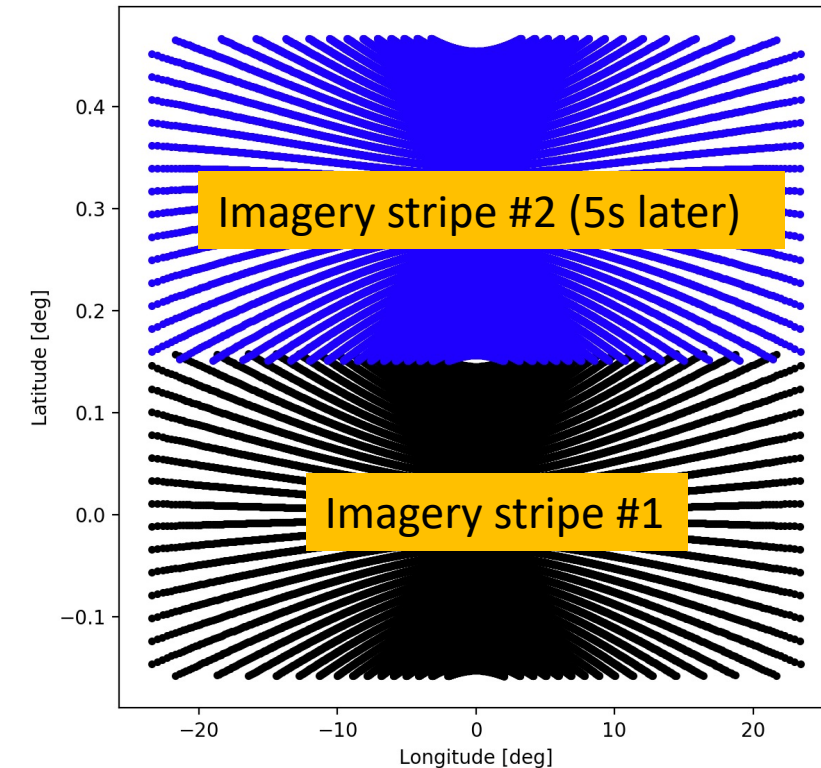


Motivation

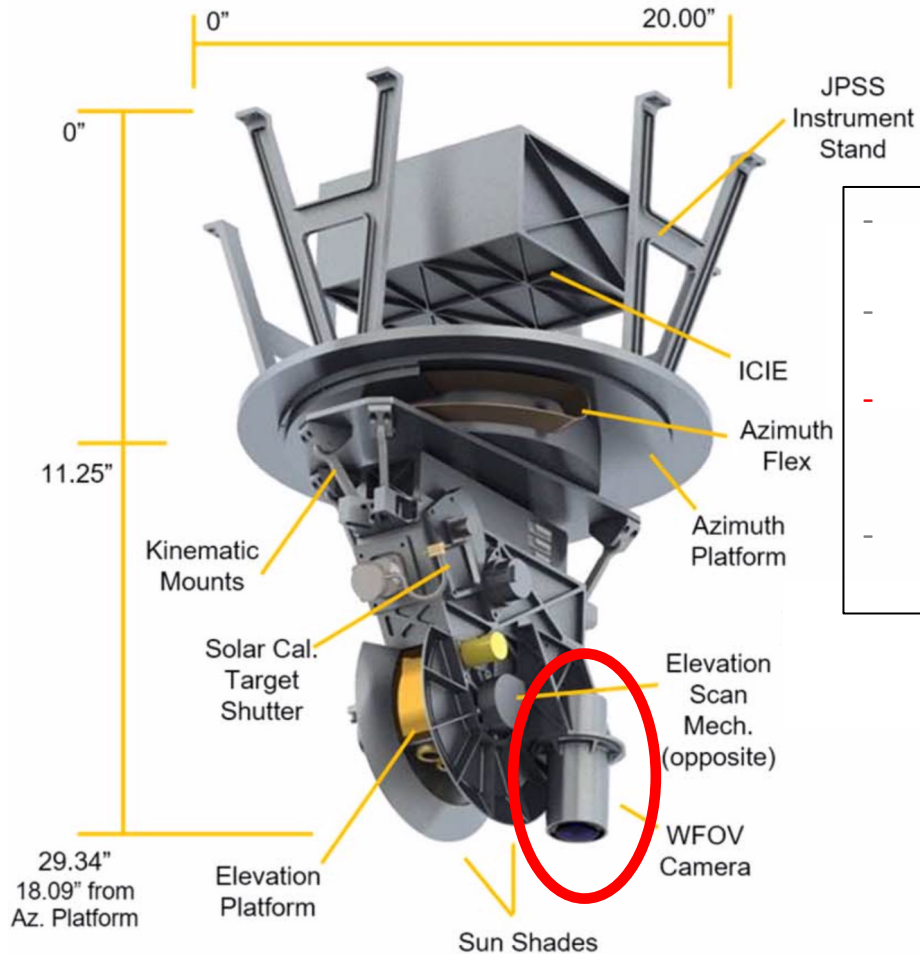


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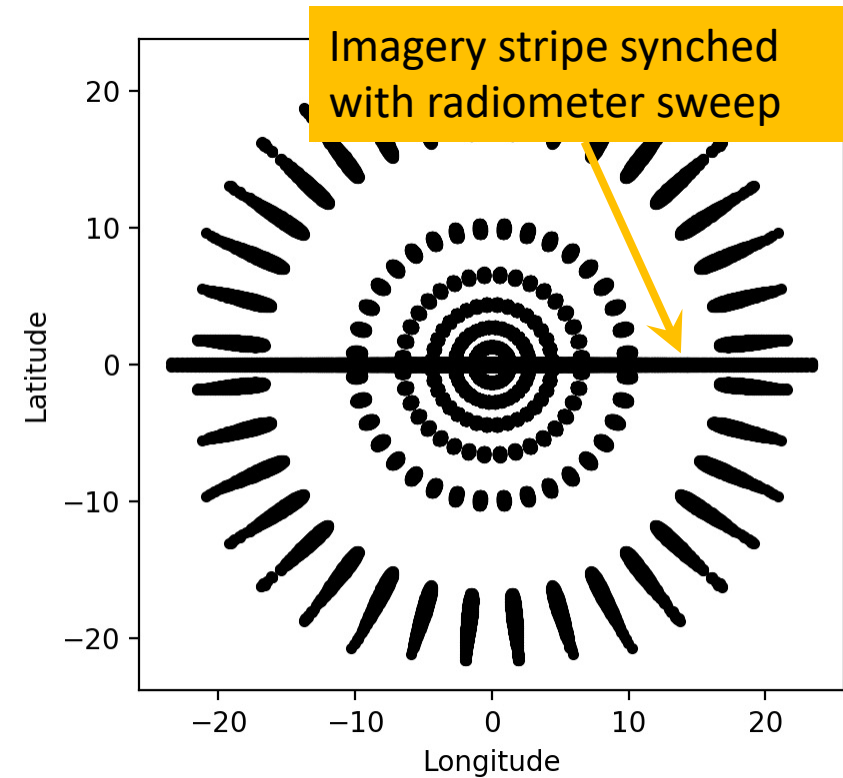


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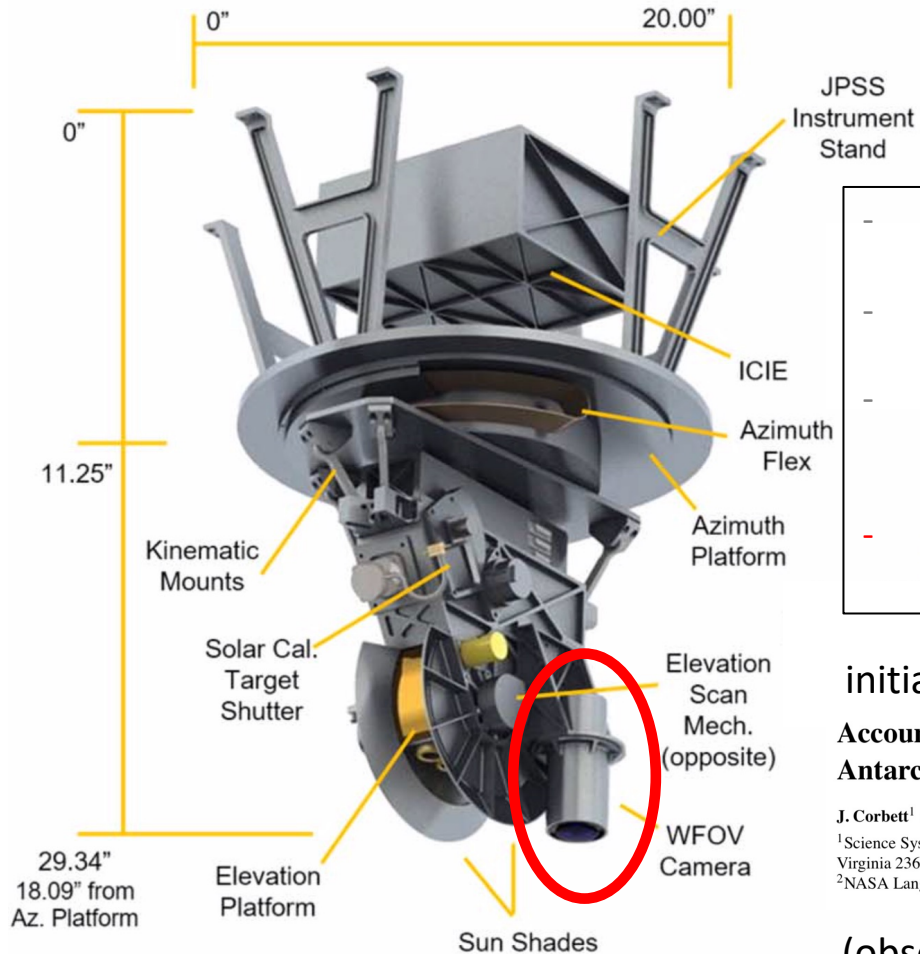
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initial motivation:

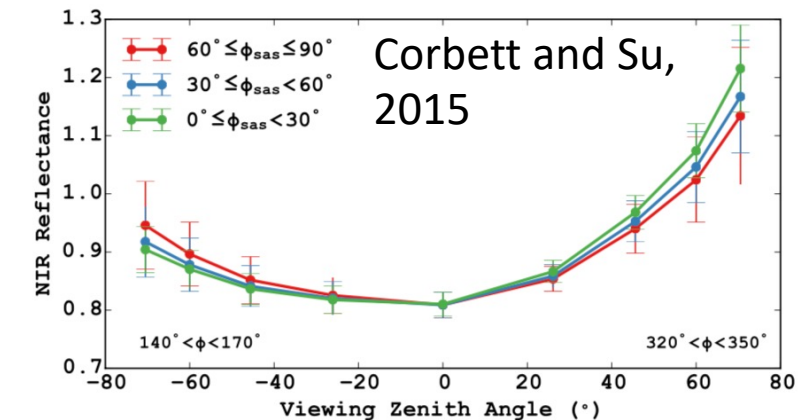
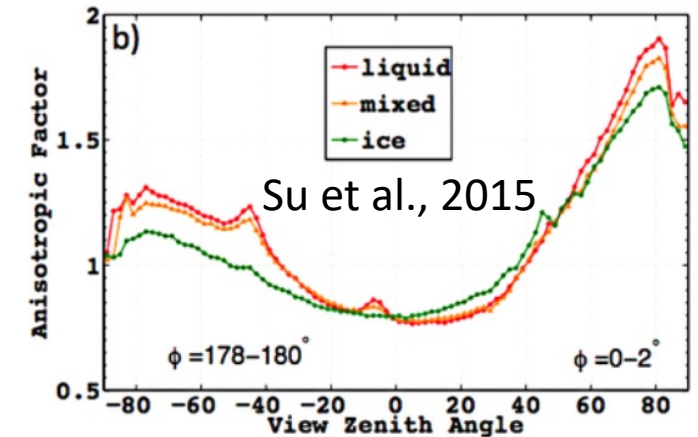
Accounting for the effects of sastrugi in the CERES clear-sky Antarctic shortwave angular distribution models

J. Corbett¹ and W. Su²

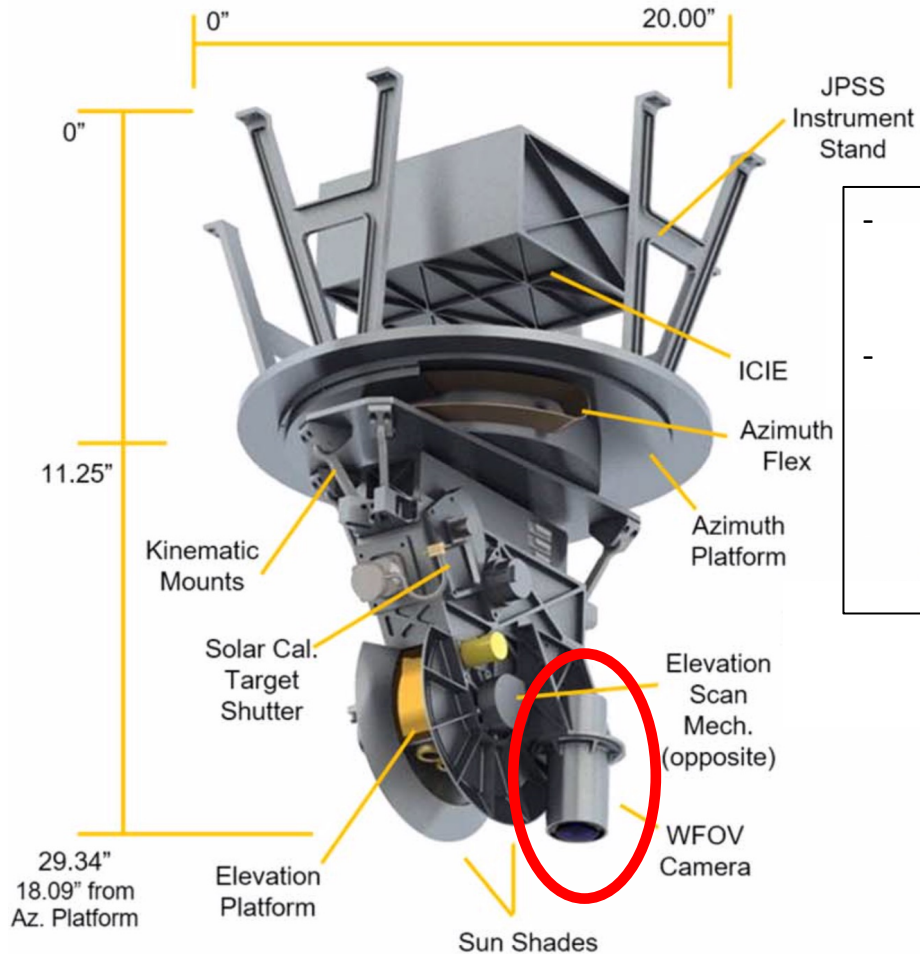
¹ Science Systems and Applications, Inc., NASA Langley Research Center, Mail Stop 420, Hampton, Virginia 23681-2199, USA

² NASA Langley Research Center, Mail Stop 420, Hampton, Virginia 23681-2199, USA

(observational instead of RT approach)



Motivation

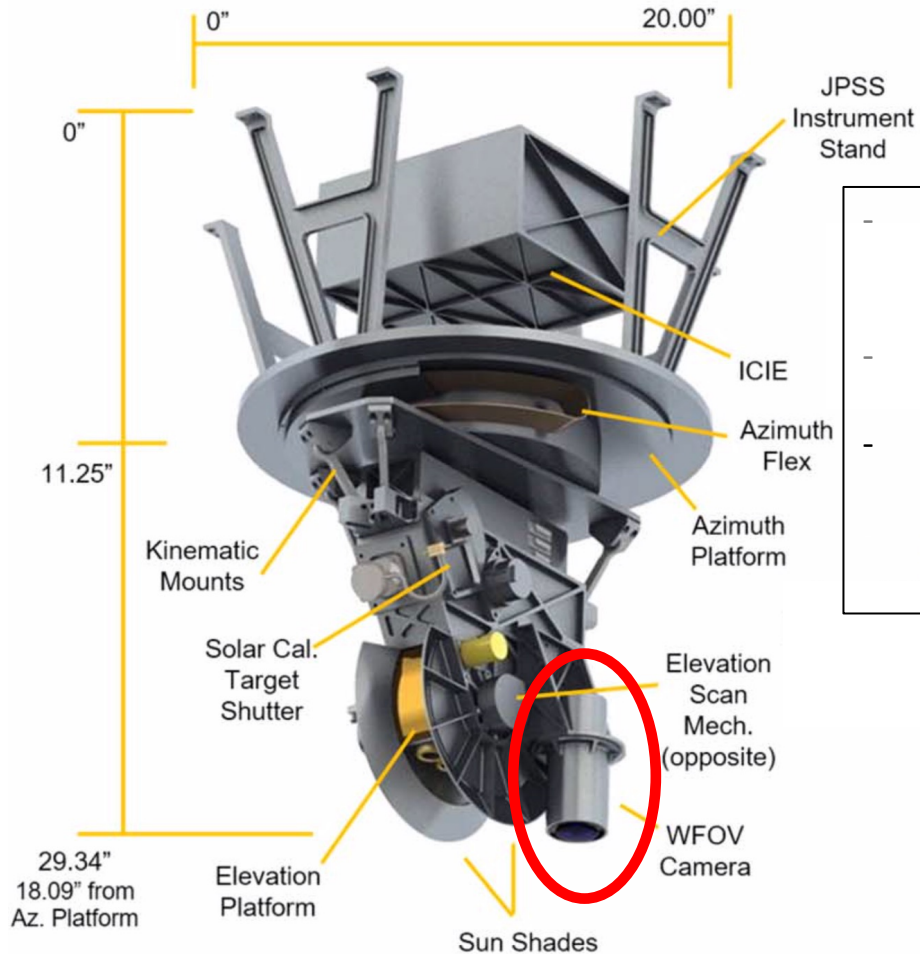


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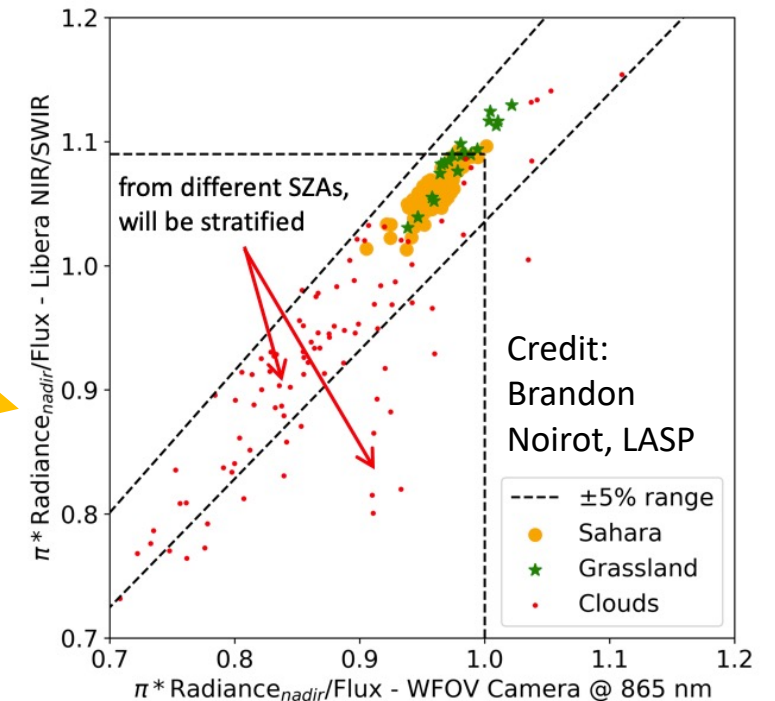
- Idea: use camera as a proxy for partial band angular distribution → get ADM faster than in RAP mode; [see Jake's talk on the details](#)
- For now, we rely on simple ERBE ADM approach to build this up

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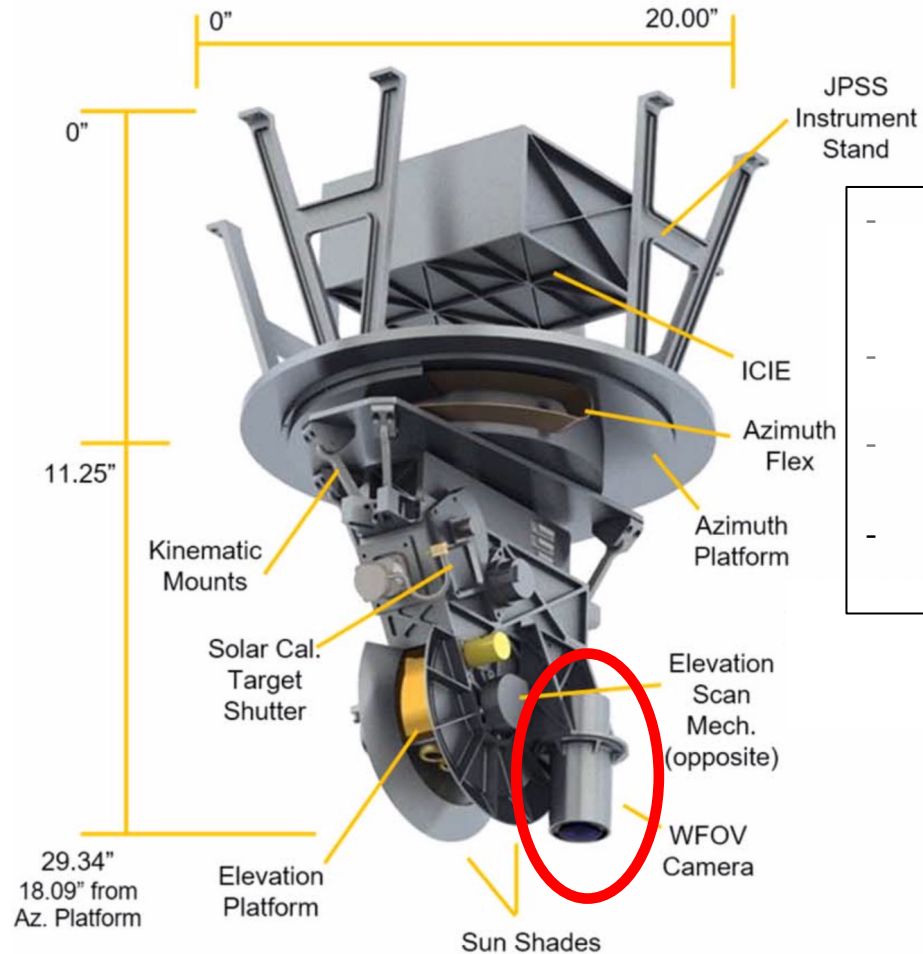
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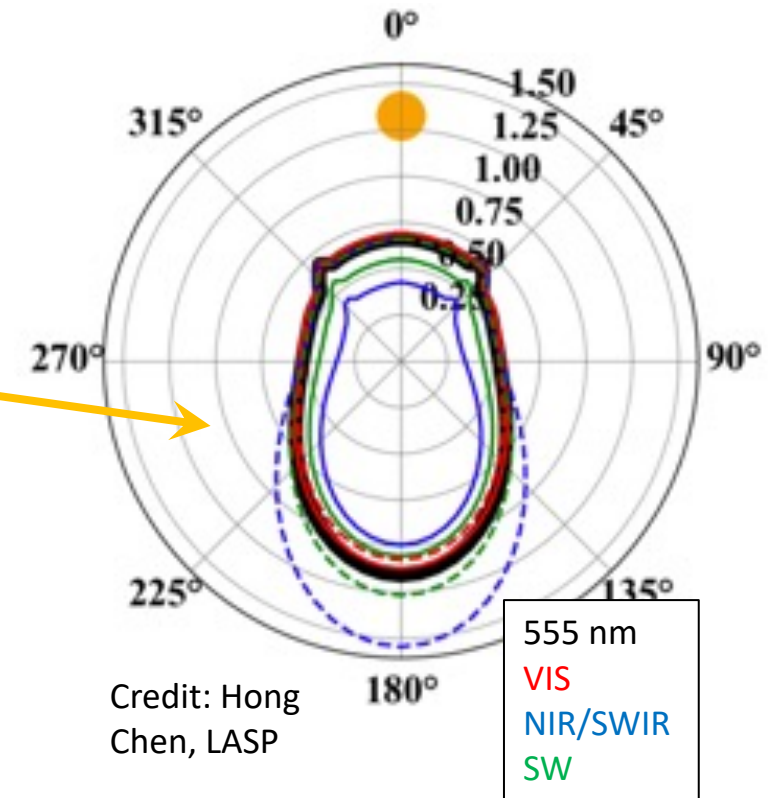


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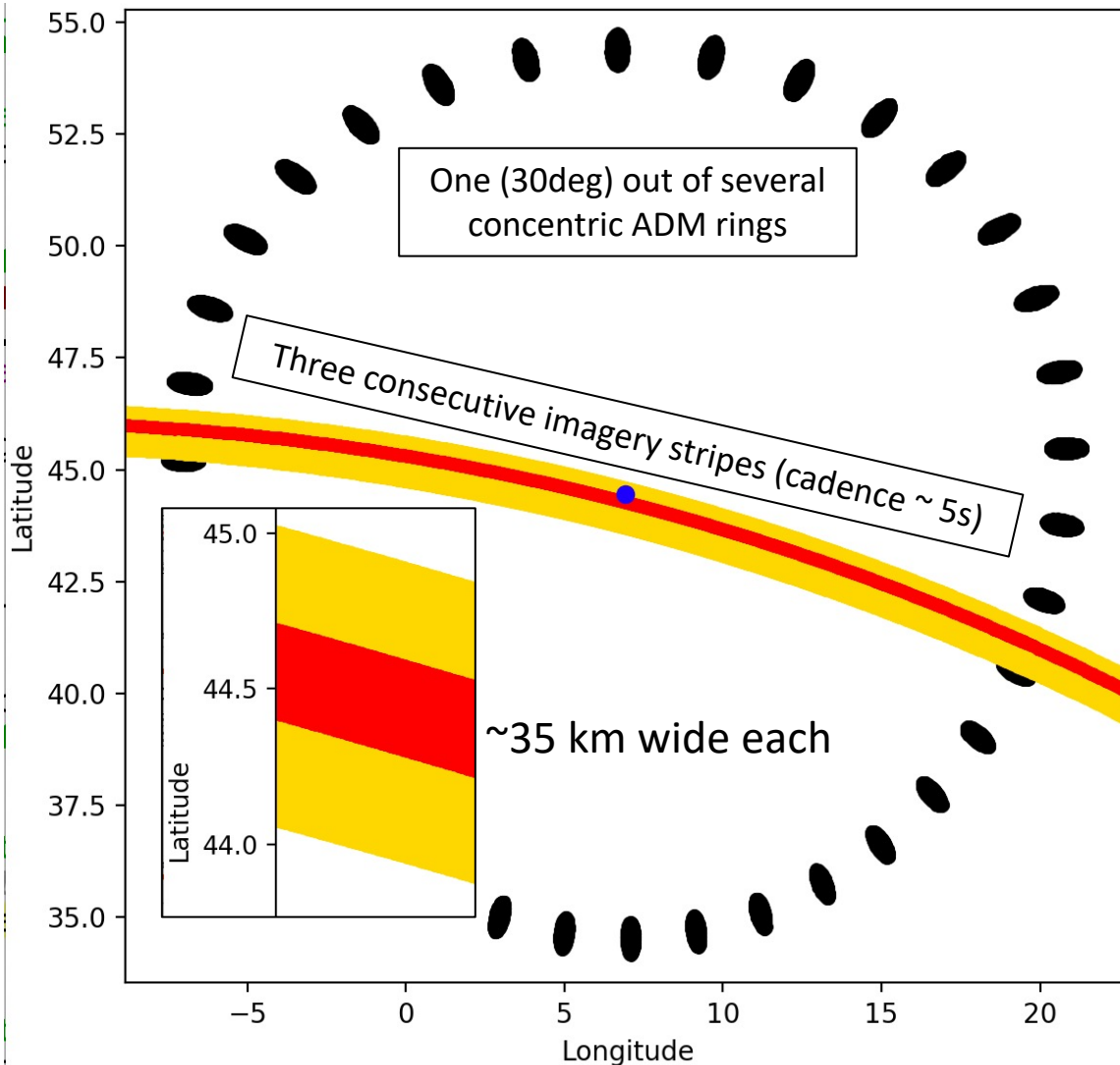


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- Radiance at 555 nm serves as angular proxy for VIS band (RT calculations)

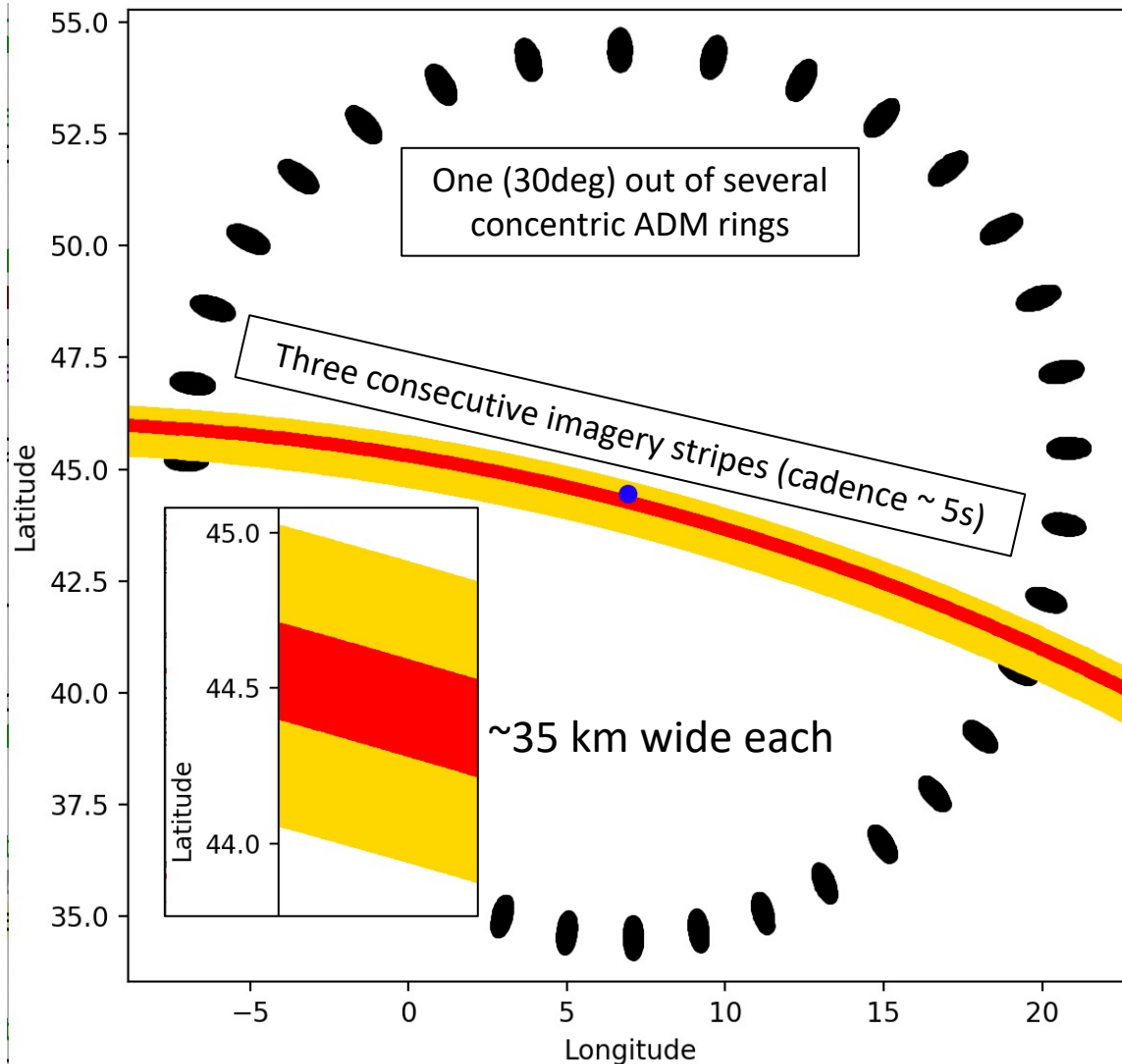


General Concept



- 2048 x 2048 pixel hemispheric camera
- 555 nm radiances (VIIRS M4: 0.545 - 0.565 micron)
- 2 overlapping dynamic range settings from 2 integration times for optimal SNR depending on pixel-level radiances
- Sub-sample central imagery stripe + ADM rings at 5 s cadence (we are data-rate limited)
- Solar shroud at ~62deg (limb) to limit degradation of optics

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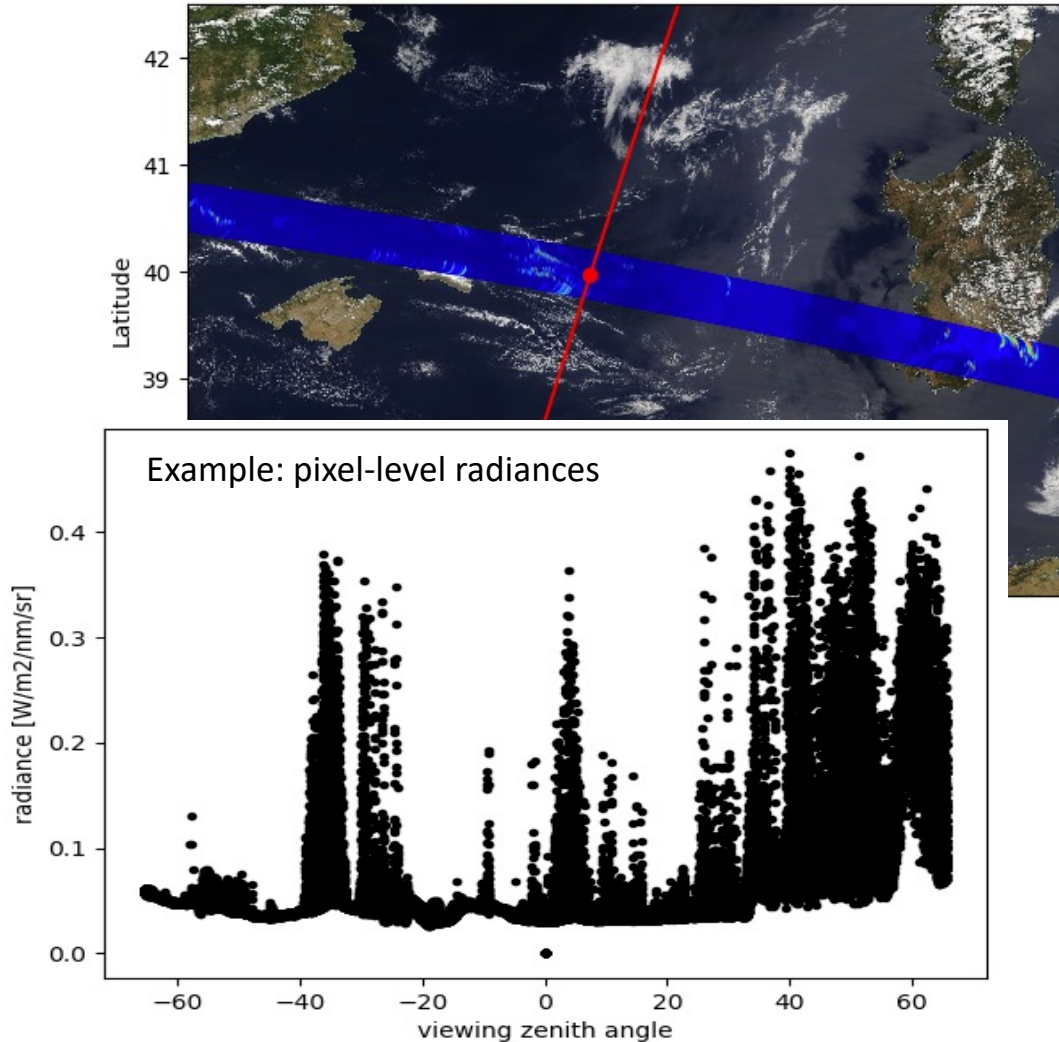


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Products:

- Cloud fraction based on adaptive radiance thresholding @ Libera PSF
- Pixel-level radiances from imagery stripe + ADM
- Libera-(PSF)-aggregated radiances along with collocated CF
 - across-track: at Libera radiometer footprints
 - along-track: CF product is translated to fw/aft angles and/or obtained directly from ADM samples

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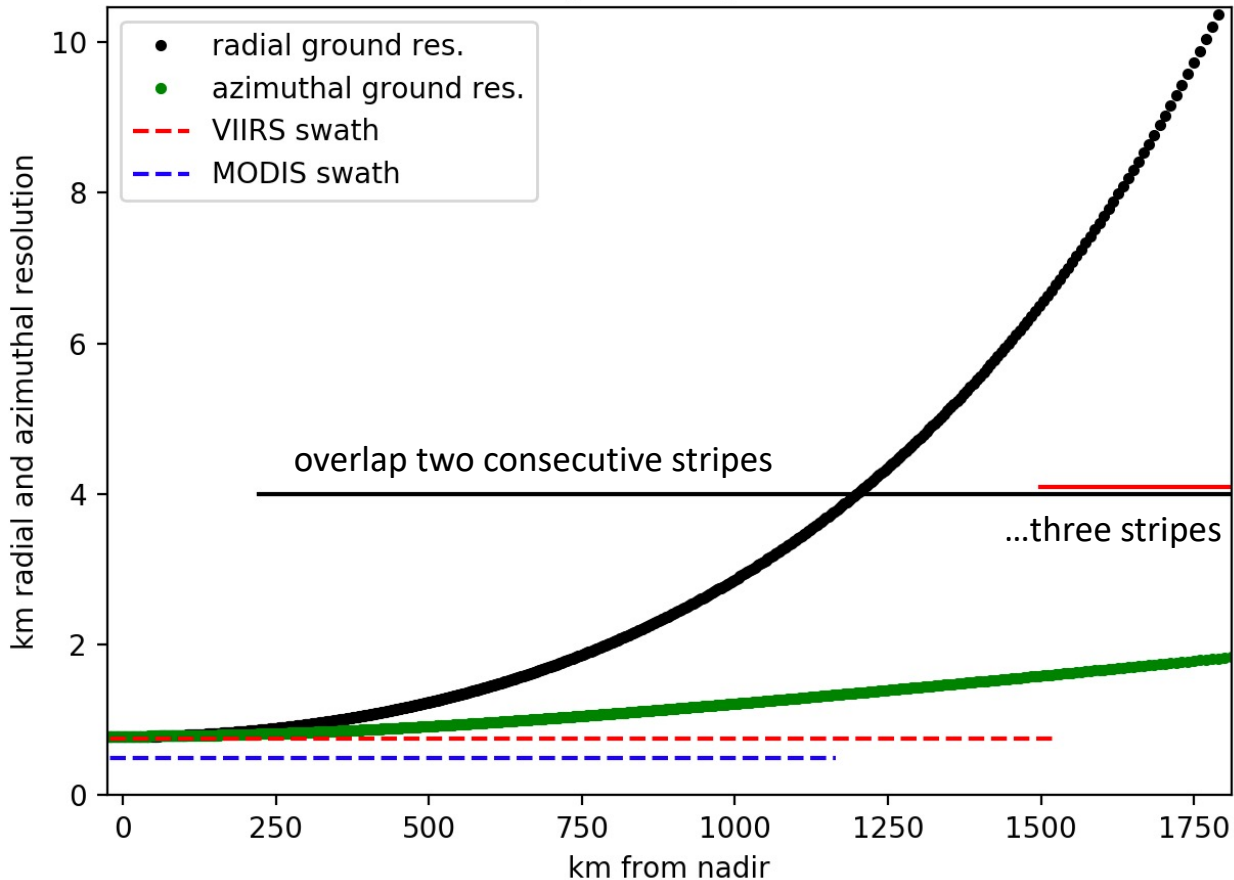


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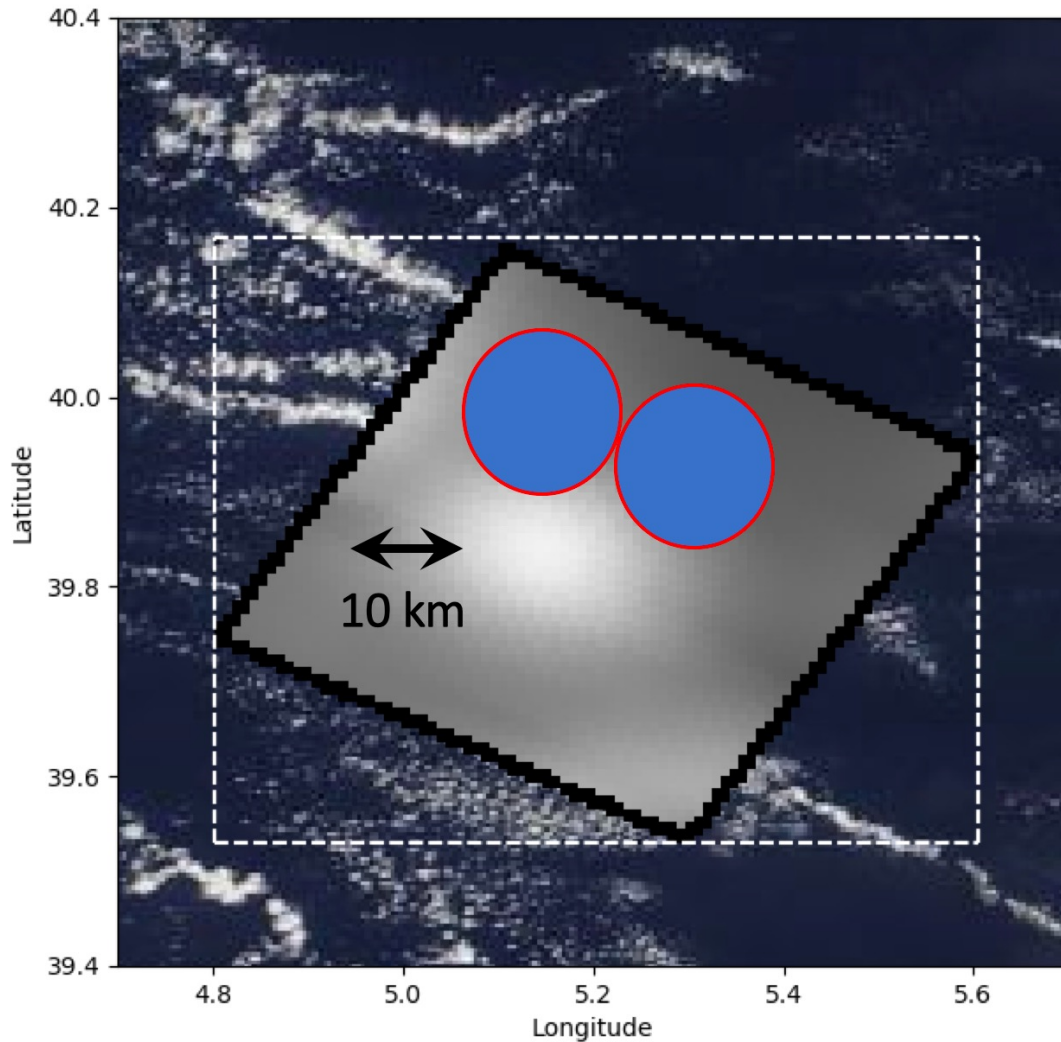


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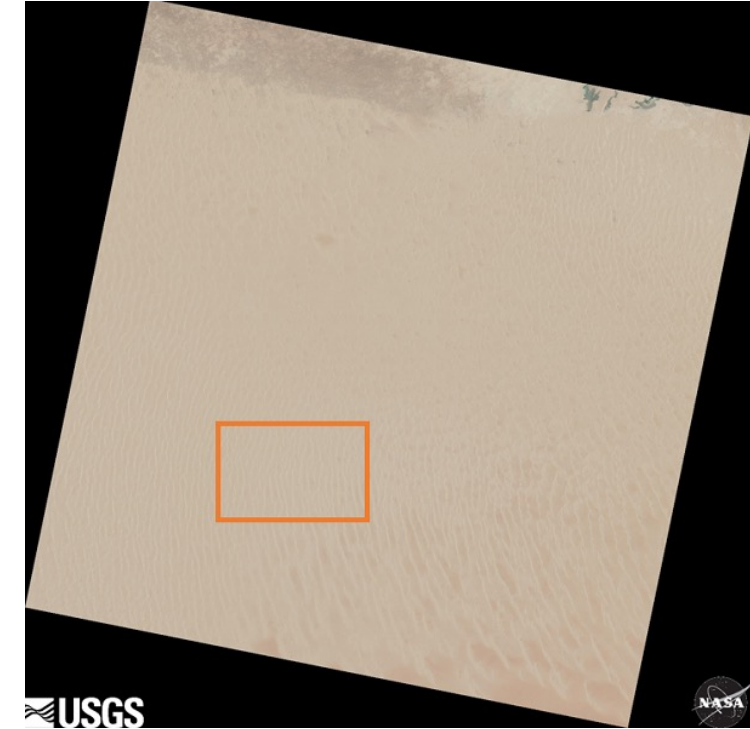
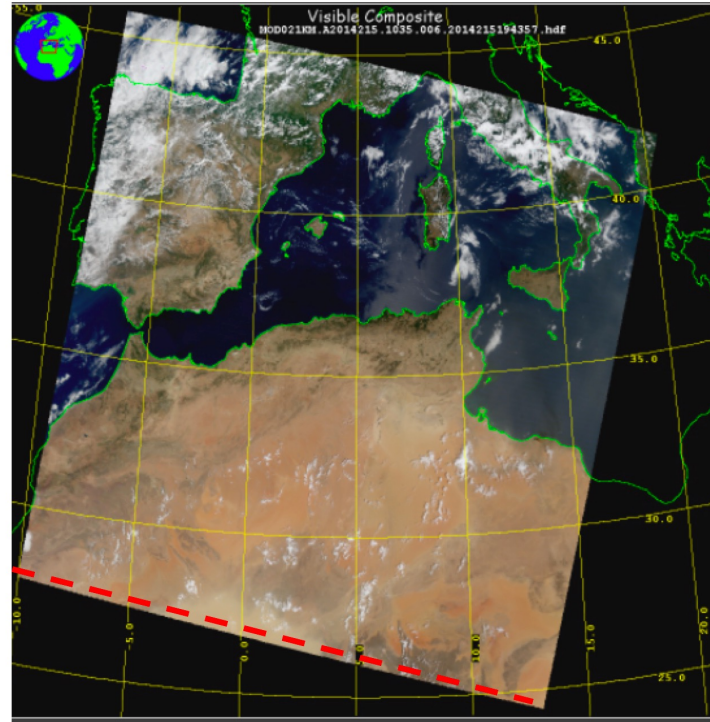
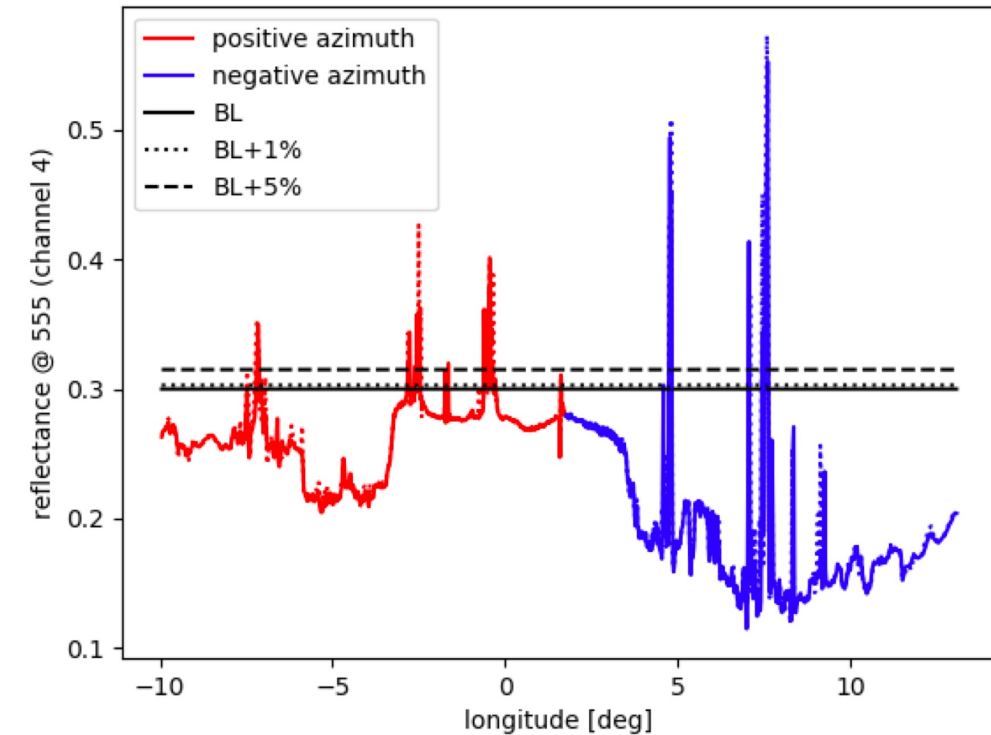
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Validation/Calibration

Cross-track: Routine regression of **Libera radiances vs. camera radiances vs. VIIRS radiances** for numerous scene types
→ check if 555 nm channel does provide proxy measurement for Libera split-solar channel angular radiance distribution
→ Enables vicarious calibration at pseudo-invariant calibration sites



Validation/Calibration

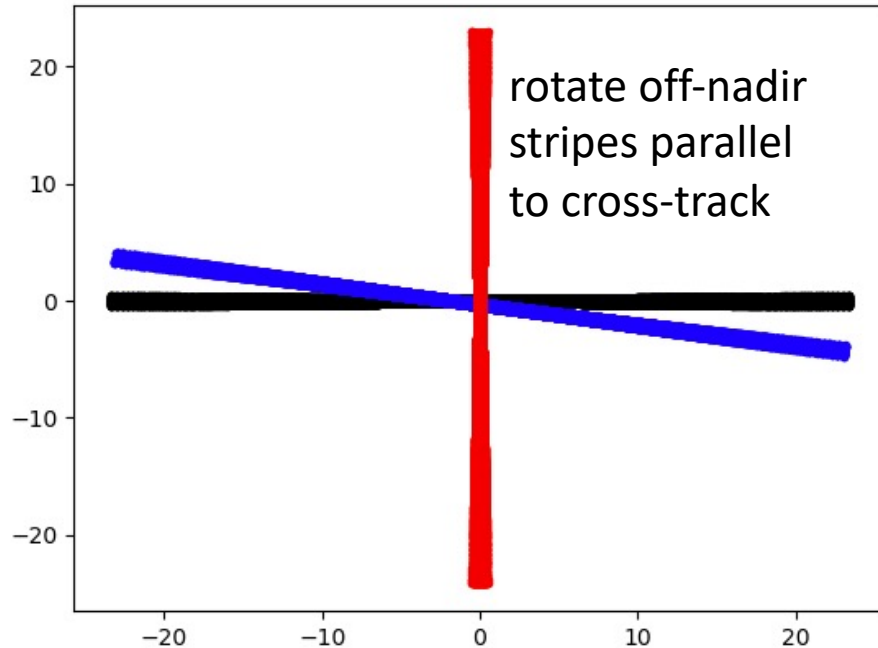
Non Cross track

Pseudo-RAP maneuvers:

- Regression of non cross-track radiances for selected scene types
- Azimuthal flat-fielding over stable targets using VIIRS M4

Lunar maneuvers:

- Use limb observations to fill in flat-fielding at periphery of FOV
- Potential use of the limb/moon for radiometric calibration



Performance Attributes*:

- Target radiometric accuracy 5%
- Target uniformity 1.5%

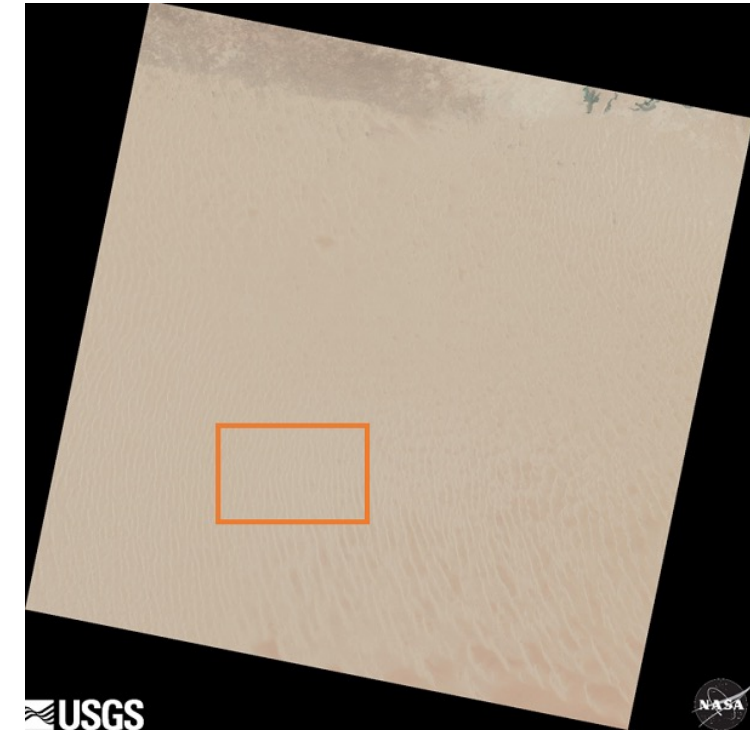
Ground Characterization

- At Ball/LASP

On-orbit

- No onboard calibrator in “demo mode”
- Rely on vicarious methods (for now)

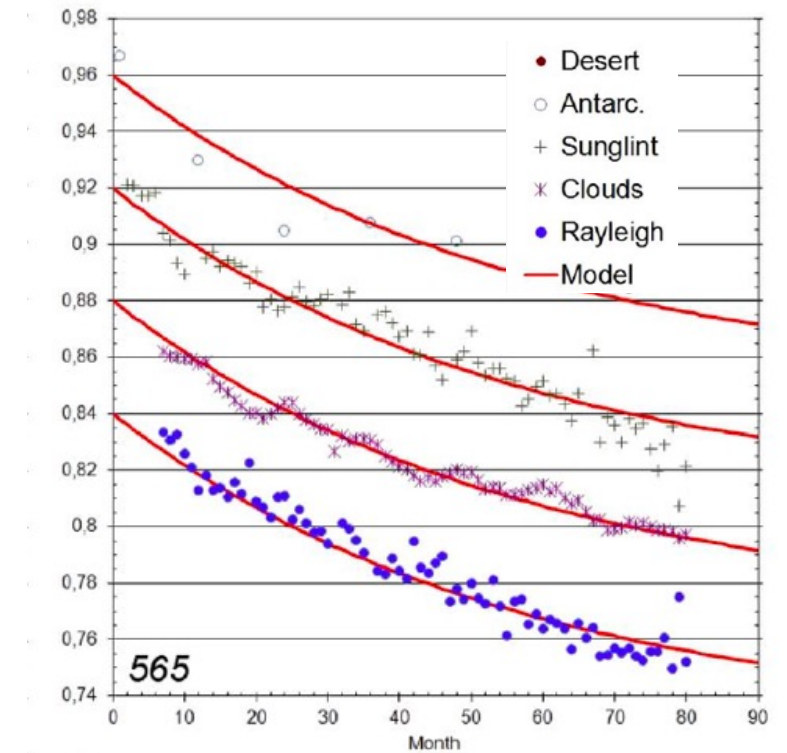
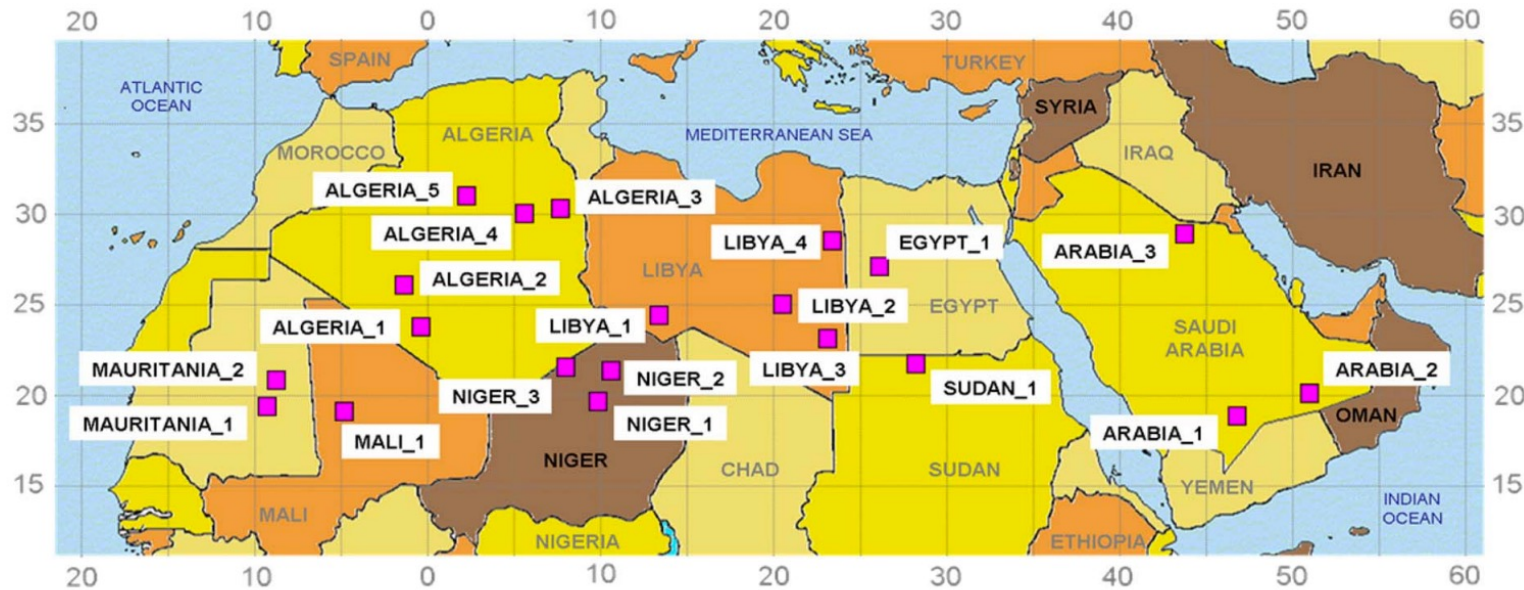
Currently flowing top-level requirements to
component-level requirements + approach



* also: filter spectral response; image distortion; detector linearity; dark current (T); straylight

Stability

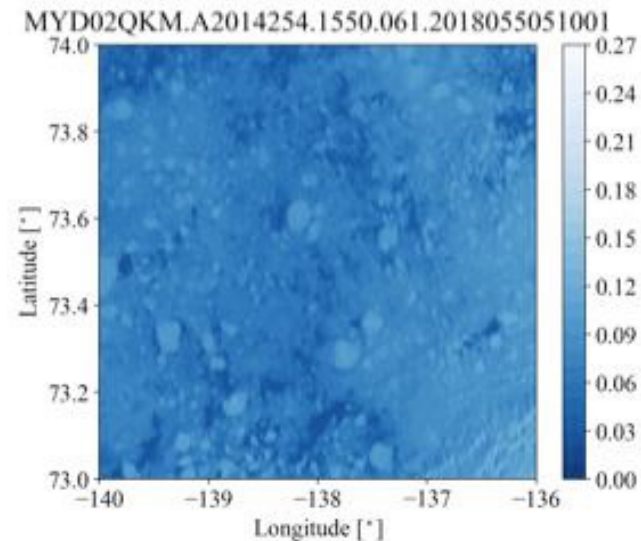
Ensure stability over time via vicarious calibration methods:
Bacour et al, 2019



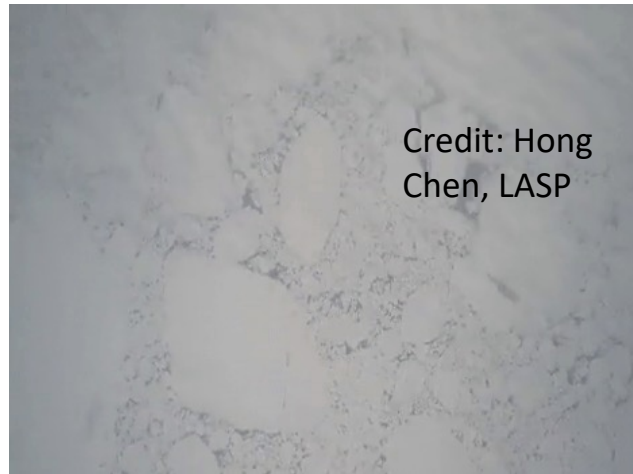
New Remote Sensing Avenues

Over bright surfaces, the adaptive thresholding method for cloud detection based on single band will fail (will nominally rely on VIIRS), **but...**

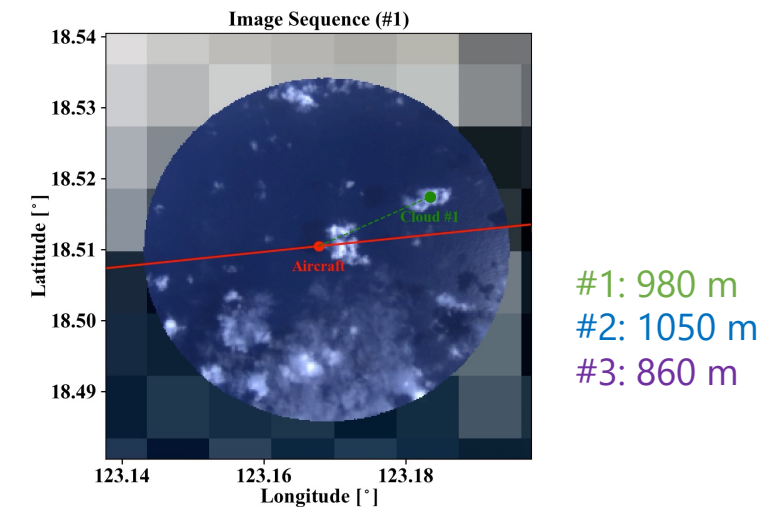
...we are working on feature tracking algorithms which detect clouds based on stereoscopic methods
...may also be able to determine cloud top height



Multi-overpass imagery of clouds and ice floes during ARISE → ARCSIX



Multi-angle airborne camera imagery during ARISE 2014



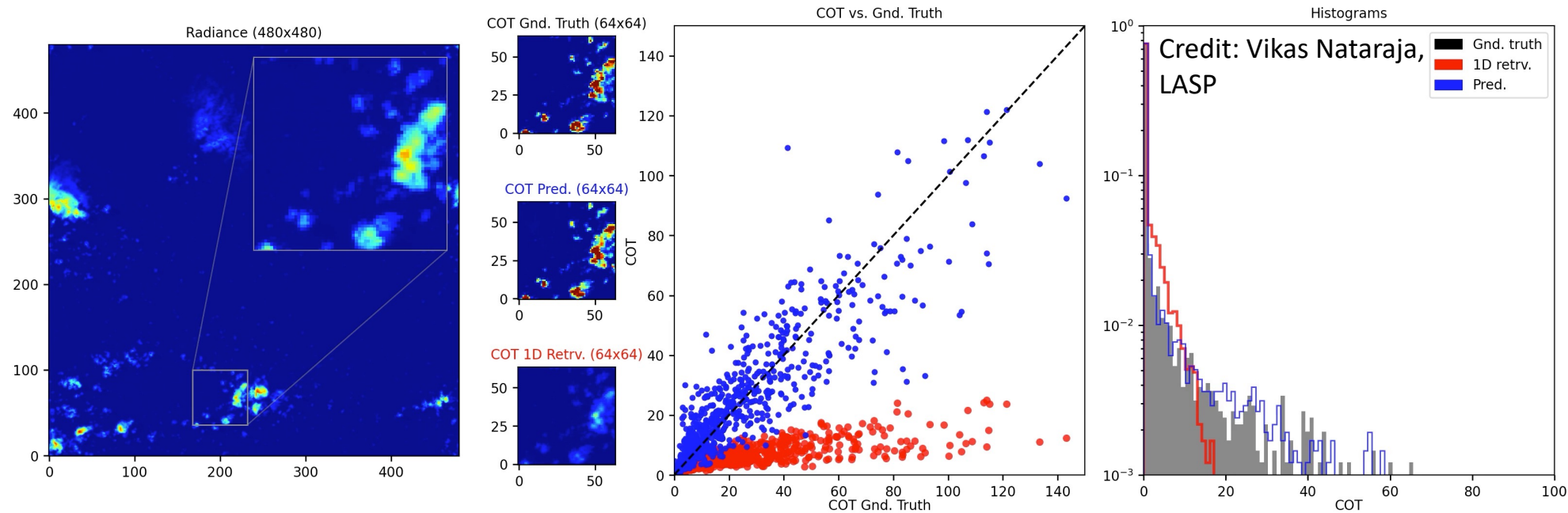
Multi-angle airborne camera imagery during CAMP²Ex 2019

New Remote Sensing Avenues

From one single channel, we are not expecting retrievals of cloud optical properties, **but...**

...we developed a Convolutional Neural Network retrieval for COD (so far, for water clouds, Nataraja et al., 2021), which detects/corrects for 3D effects

...various multi-angle/multi-pixel approaches* will be available for AtmOS → piggy-back



- Multi-angle algorithms (Shi et al., 2008); feature tracking (Kölling et al., 2019) → stereo height
- Multi-pixel algorithms (Jafariserajehlou et al., 2019) → thin cloud detection even over bright surfaces
- Convolutional neural networks (Nataraja et al., 2021) + cloud tomography (Levis et al., 2020)

CERES/Libera STM, May 11-13, 2021

Summary

Camera accelerates split-solar channel ADM aggregation (use 555 nm as proxy for Libera-VIS in off-cross-track direction)

New Remote Sensing Avenues that rely on limb-to-limb angular radiance measurements

Full definition of sampling and attributes still ongoing → feedback in the process and ideas are solicited

